

4-13-2023

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Dong, W., Fritts, D., Liu, A. Z., Liu, H., & Snively, J. (2023). Accelerating Atmospheric Gravity Wave Simulations using Machine Learning: Kelvin-Helmholtz Instability and Mountain Wave Sources Driving Gravity Wave Breaking and Secondary Gravity Wave Generation. *Research Square*, (). <https://doi.org/10.21203/rs.3.rs-2790920/v2>

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Article

Keywords: Machine Learning, Kelvin-Helmholtz instability, Gravity Wave

Posted Date: April 13th, 2023

DOI: <https://doi.org/10.21203/rs.3.rs-2790920/v2>

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Additional Declarations: There is **NO** Competing Interest.

1 **Accelerating Atmospheric Gravity Wave Simulations using Machine**
2 **Learning: Kelvin-Helmholtz Instability and Mountain Wave Sources Driving**
3 **Gravity Wave Breaking and Secondary Gravity Wave Generation**

4
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7
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15 **Key Words:** Machine Learning, Kelvin-Helmholtz instability, Gravity Wave

16 **Key Points:**

- 17 1. A machine learning model (CAMNet) tailored for nonlinear gravity wave
18 simulations is developed.
- 19 2. CAMNet can achieve a several order-of-magnitude acceleration relative to
20 physics-based model without sacrificing accuracy.
- 21 3. CAMNet opens a new window to improve the parameterization of primary and
22 secondary GWs in the global atmospheric models.

23 Abstract

24 Gravity waves (GWs) and their associated multi-scale dynamics are known to play
25 fundamental roles in energy and momentum transport and deposition processes
26 throughout the atmosphere. We describe an initial, two-dimensional (2-D), machine
27 learning model – the Compressible Atmosphere Model Network (CAMNet) - intended
28 as a first step toward a more general, three-dimensional, highly-efficient, model for
29 applications to nonlinear GW dynamics description. CAMNet employs a physics-
30 informed neural operator to dramatically accelerate GW and secondary GW (SGW)
31 simulations applied to two GW sources to date. CAMNet is trained on high-resolution
32 simulations by the state-of-the-art model Complex Geometry Compressible
33 Atmosphere Model (CGCAM). Two initial applications to a Kelvin-Helmholtz
34 instability source and mountain wave generation, propagation, breaking, and SGW
35 generation in two wind environments are described here. Results show that CAMNet
36 can capture the key 2-D dynamics modeled by CGCAM with high precision. Spectral
37 characteristics of primary and SGWs estimated by CAMNet agree well with those
38 from CGCAM. Our results show that CAMNet can achieve a several order-of-
39 magnitude acceleration relative to CGCAM without sacrificing accuracy and suggests
40 a potential for machine learning to enable efficient and accurate descriptions of
41 primary and secondary GWs in global atmospheric models.

42 Plain Language Summary

43 Atmospheric gravity waves (GWs) are well described by the Navier-Stokes equations,
44 but solving these equations including small scale remains daunting, limited by the
45 very high computational cost of resolving the smallest spatial-temporal features in a
46 global context. To address this challenge, we developed a machine learning model
47 called CAMNet. Our model demonstrates that neural networks can be trained on
48 high-resolution compressible atmospheric model data and then used to simulate
49 gravity wave evolution. Importantly, initial results show that using such trained
50 model can achieve computational savings of >1000 times compared to a physics-based

51 simulation while still achieve highly accurate results. These findings are exciting, as
52 they suggest that CAMNet can overcome the limitations of current GW
53 parameterizations and provide a promising avenue for studying the effects of sub-
54 grid-scale processes in atmospheric science and properly incorporating them in global
55 models. The development of CAMNet opens up major new opportunities for improving
56 effective model resolution, accuracy, and efficiency.

57 1. Introduction

58 Gravity waves (GWs) play prominent roles throughout Earth's atmosphere. They are
59 generated at lower altitudes by various primary sources including airflow over
60 topography (i.e., mountain waves, MWs), convection, and jet streams. Additional
61 GWs are generated due to strong GW/mean-flow interactions described as “self-
62 acceleration” (SA) dynamics and by resonant and off-resonant wave-wave
63 interactions that can take many forms. These diverse GWs play central roles in
64 Earth's atmospheric dynamics and climate by transporting energy, pseudo-
65 momentum, and constituents over depths extending into the thermosphere (Fritts &
66 Alexander, 2003).

67 Accounting for larger- and smaller-scale GW transports and influences remains a
68 challenging problem due to the complex physics involved and the need for high-
69 resolution simulations to describe detailed responses where these are important.
70 Traditional simulation methods, such as finite difference or finite volume methods,
71 can be computationally expensive, and quantifying sub-grid scale processes has been,
72 and remains, a major challenge. Multiple parameterization schemes spanning 40
73 years have aimed to account for GW pseudo-momentum deposition for various GW
74 sources, discretely or spectrally (i.e., linearly or nonlinearly), from the surface into
75 the thermosphere (e.g., Lindzen, 1981; Holton, 1982; Palmer et al., 1986; Fritts and
76 Lu, 1993). More recent schemes have built on these earlier efforts and insights
77 (Warner and McIntyre, 1996; Hines, 1997; Alexander & Dunkerton, 1999; Yiğit et al.,

78 2008; Eckermann et al., 2015; Amemiya & Sato, 2016; Gettleman et al., 2019; Miyoshi
79 & Yiğit., 2019; Ribstein et al., 2022).

80 Importantly, all these various schemes are based on simplified, often linear or weakly
81 nonlinear, mathematical models and/or empirical relations that are significant
82 approximations having limited quantitative predictive abilities. As such, they
83 introduce significant model uncertainties and biases in predictions of middle and
84 upper atmosphere responses (Pedatella et al., 2014). Additionally, parameter settings
85 in these schemes may require adjustments for different models, model configurations,
86 and/or model resolutions.

87 Parameterizations addressing GWs that are partially resolved (the "gray zone") that
88 maintain physical consistency between the resolved and parameterized dynamics are
89 promising (Vosper, 2015; Vosper et al., 2016), but can also be challenging (Liu, 2019).
90 Such efforts attempt to represent the complex and highly nonlinear physics of GWs.
91 However, there remain many aspects of GW dynamics, e.g., SA dynamics, local
92 instabilities and breaking, multi-scale interactions, and secondary GW (SGW)
93 generation that become increasingly important at increasing altitudes, but that
94 cannot be addressed by linear theory or existing GW parameterization schemes.

95 The recent boom in hardware and software developments relevant to machine
96 learning (ML) has motivated some efforts to examine the possible benefits that ML
97 can bring to GW parameterization (e.g., Chantry et al., 2021; Espinosa et al., 2022).
98 ML methods offer several potential advantages over traditional parameterization
99 schemes. These include the following:

- 100 1) ML applications can learn complex, nonlinear relationships directly from
101 data, without the need for pre-determined equations or assumptions. This
102 makes ML methods well-suited for problems exhibiting highly nonlinear
103 dynamics;

104 2) ML can identify complex patterns and relationships in the data that may be
105 difficult to discern using traditional methods. This can lead to improved
106 accuracy and transferability of the resulting flow descriptions; and

107 3) ML has the potential to dramatically reduce the computational cost
108 associated with descriptions of nonlinear GW dynamics by replacing
109 traditional parameterizations with highly-efficient, data-driven models.

110 Such methods have been shown to yield significant benefits in several applications to
111 date. Chantry et al. (2021) trained a neural network on an upgraded version of an
112 existing parameterization scheme that yielded improved results describing GW drag
113 in a numerical weather prediction (NWP) system. In another study, Espinosa et al.
114 (2022) developed an artificial neural network to emulate the pseudo-momentum
115 forcing described in a traditional GW parameterization in an idealized climate model.
116 By coupling the climate model with their ML-based GW parameterization, they were
117 able to accurately reproduce the quasi-biennial oscillation, a well-known atmospheric
118 phenomenon. However, these ML-based GW parameterizations rely on traditional
119 GW parameterizations, hence inherit their assumptions and simplifications. They
120 also remain unable to represent the true, highly nonlinear GW dynamics.

121 As mentioned earlier, the dynamics of GWs are governed by the Navier-Stokes
122 equations. In recent years, several ML-based solvers for partial differential equations
123 (PDEs) have been proposed to approximate or improve various numerical methods.
124 The most explored of these can be divided into two categories: physics-informed
125 neural networks (PINN, e.g., Maziar et al., 2019; Wandel et al., 2022) and neural
126 operators (NOs, e.g., Lu et al., 2019; Li et al., 2020; Xiong et al., 2023). PINN uses a
127 neural network as the solution function and optimizes a loss function to minimize
128 violation of the given equation. However, it experiences difficulties in propagating
129 information from initial or boundary conditions to unseen parts of the interior and to
130 future times. NOs are better suited for solving PDEs and have been successfully used
131 in flow prediction (e.g., Lu et al., 2019; Li et al., 2020; Xiong et al., 2023). However,

132 they require large volumes of simulation data. Recently, physics-informed neural
133 operators, e.g., both the physics-informed Deep Operator Network proposed by
134 Goswami et al(2022) and physics-informed Fourier neural operator proposed by Li et
135 al (2022) employ both data and physics losses on operator learning to overcome the
136 shortcomings of purely PINN or data-driven learning. Our ML approach is also based
137 on this physics-informed neural operator.

138 While MLs offer potential advantages for GW simulation, there are also challenges
139 that must be addressed. The primary challenge is the limited availability of high-
140 quality data for training and validation. To address this need, we utilize the Complex
141 Geometry Compressible Atmospheric Model (CGCAM), a finite volume model that
142 has been used extensively to study GW dynamics and their instabilities in the
143 Mesosphere and Lower Thermosphere (MLT) at very high resolution, see, e.g., Dong
144 et al., 2020, 2021,2022,2023; Fritts et al., 2020, 2021,2022a, 2022b; Lund et al., 2021).
145 CGCAM is capable of capturing highly nonlinear GW and GW-related dynamics
146 yielding high-fidelity GW training datasets for ML-based approaches.

147 Inspired by previous research in this field, this study investigates the application of
148 ML algorithms for simulating GW dynamics. Our focus is on the potential of ML to
149 improve the efficiency of GW simulations while capturing their highly nonlinear
150 dynamics with high fidelity, specifically including instabilities, breaking, and SGW
151 generation. Our approach employs the Compressible Atmosphere Model Network
152 (CAMNet) model based on CGCAM that solves the compressible Navier-Stokes
153 equations in the Complex Geometry Compressible Atmosphere Model (CGCAM). To
154 explore the performance of CAMNet, we train the model using single-channel inputs
155 instead of multi-channel inputs as explored by Pathak et al (2022), with a focus on
156 improving training efficiency. We expect this study to provide valuable insights into
157 the potential benefits of ML in modeling GW and instability dynamics and to inform
158 future research in this area. When optimized, this method will be extended to three-
159 dimensions in order to achieve much more efficient descriptions of GW effects in
160 global models while achieving sufficient accuracy.

161 **2. Method**

162 **2.1 Complex Geometry Compressible Atmospheric Model (CGCAM)**

163 The Complex Geometry Compressible Atmosphere Model (CGCAM) solves the
 164 two/three-dimensional (2/3-D) nonlinear and compressible Navier-Stokes equations
 165 written in strong conservation law (flux) form as follows:

166
$$\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u_j)}{\partial x_j} = 0 \quad (1)$$

167
$$\frac{\partial(\rho u_i)}{\partial t} + \frac{\partial(\rho u_i u_j)}{\partial x_j} = -\frac{\partial p}{\partial x_i} - \rho g \delta_{i3} + \frac{\partial \sigma_{ij}}{\partial x_j} \quad (2)$$

168
$$\frac{\partial(\rho E)}{\partial t} + \frac{\partial[(\rho E + p)u_j]}{\partial x_j} = -\rho g u_3 + \frac{\partial(u_i \sigma_{ij})}{\partial x_j} - \frac{\partial q_j}{\partial x_j} \quad (3)$$

169 where σ_{ij} and q_j are the viscous stress and thermal conduction, respectively, defined
 170 as

171
$$\sigma_{ij} = \mu \left[\left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) - \frac{2}{3} \left(\frac{\partial u_k}{\partial x_k} \right) \delta_{ij} \right] \quad (4)$$

172
$$q_j = -\kappa \frac{\partial T}{\partial x_j} \quad (5)$$

173 and where μ is the dynamic viscosity, κ is the thermal conductivity, δ_{ij} is the
 174 Kronecker delta, ρ is density, and g is the gravitational acceleration. μ and κ depend
 175 on the temperature through Sutherland's Law (White, 1974).

176 The solution variables are ρ , the momentum per unit volume, ρu_i or $(\rho u, \rho v, \rho w)$, and
 177 the total energy $E = e + u_k u_k / 2 = c_v T + u_k u_k / 2$, with velocity components
 178 (u_i, u_j, u_k) along (x, y, z) . Also $c_v = \frac{R}{\gamma - 1}$ is the specific heat at constant volume and T is
 179 the temperature. The compressible equation set is discretized using a second-order

180 finite-volume scheme identical to the method discussed by Felten & Lund (2006).
181 Time advancement is achieved via a third-order accurate Runge-Kutta scheme.
182 Additional details for CGCAM are provided by Dong et al. (2020) and Lund et al.
183 (2020). CGCAM has been successfully used in various studies on GW generation,
184 breaking, and SGW generation (e.g., Dong et al., 2020,2021,2022; Fritts et al.,
185 2020;2021; 2022a; 2022b; Lund et al., 2021).

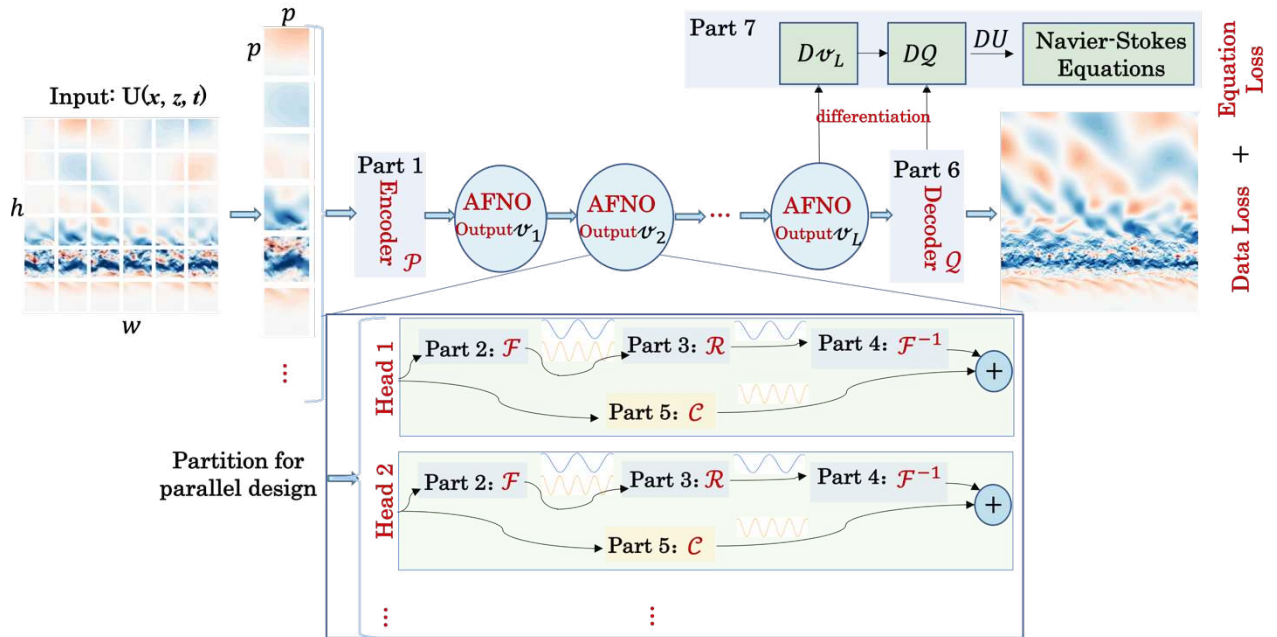
186 **2.2 Compressible Atmospheric Model Network (CAMNet)**

187 CAMNet is a hybrid machine learning model that combines data-driven and physics-
188 informed approaches. It is based on the Adaptive Fourier Neural Operator (AFNO)
189 proposed by Guibas et al., 2021. AFNO is a Fourier transform-based token-mixing
190 scheme with a vision transformer backbone (Dosovitskiy et al., 2020). AFNO is based
191 on the Fourier neural operator (FNO) that learns in a resolution-invariant manner
192 and has shown success in modeling challenging partial differential equations such as
193 fluid dynamics (Li et al., 2020). The Fourier architecture of AFNO applies a fast
194 Fourier Transform (FFT) to the data and applies its fully connected layers in Fourier
195 space before performing an inverse FFT back to real space. Moreover, the Fourier
196 architecture has been demonstrated the ability to perform zero-shot super-resolution,
197 predicting on higher-resolution data having only seen low resolution data. The
198 introduction of vision transformer enables it to model long-range dependencies well
199 and yields a state-of-the-art high-resolution model that resolves fine-grained features
200 and scales well with resolution and size of dataset. AFNO enables training high-
201 fidelity data-driven models as truly unprecedented resolution (Pathak et al., 2022).

202 The power of AFNO stems from its ability to combine linear integral operators,
203 implemented through the Fourier transform, with non-linear activation functions,
204 enabling it to learn highly non-linear operators. This is similar to standard Multi-
205 Layer Perceptron (MLP) and Convolutional Neural Network (CNN), where linear
206 multiplications are combined with non-linear activations to learn highly non-linear
207 functions. Although AFNO truncates higher frequency modes in the Fourier layer, Li

208 et al. (2020) argue that the entire operator can still approximate functions with the
 209 full frequency range, due to the function being represented in a high-dimensional
 210 channel space. The non-linear decoder network then recovers the higher frequency
 211 modes when projecting back to the desired dimension. In our case of predicting multi-
 212 scale GW dynamics, the Fourier layer truncation of high-frequency information
 213 resulted in poor small-scale structure prediction. To address this, we added a
 214 convolutional layer, which is able to amplify high-frequency components and
 215 complement the information truncated by the Fourier layer.

216 Additionally, CAMNet further extends the AFNO architecture by incorporating
 217 physics information from the Navier-Stokes equations, which govern the GW
 218 dynamics. These equations are used to create a loss function that captures the
 219 violation of these laws, and Fourier derivatives (Li et al., 2021) are used to compute
 220 the derivatives for the physics constraints, as the automatic differentiation in
 221 PyTorch is very memory-intensive for this type of architecture. The physical
 222 constraints reduce the demand for training datasets and improve the generalization
 223 and physical validity of CAMNet learning compared to purely data-driven methods.



224

225 **Figure 1.** CAMNet architecture that utilizes the modified Adaptive Fourier Neural
 226 Operator (AFNO) and follows a patch-based approach. The input frame is divided
 227 into a $h \times w$ grid of patches, each of size $p \times p$, and encoded in a higher dimensional
 228 space with position embedding is added to form a sequence of tokens. These tokens
 229 are then mixed spatially using AFNO, which is repeated for L layers, and then a
 230 decoder reconstructs the patches for the next frame. A single AFNO layer is
 231 composed of multiple heads for parallel processing. Parts 2-4 of the layer applies the
 232 Fourier transform \mathcal{F} to the input, followed by a linear transform \mathcal{R} that acts on the
 233 lower-frequency Fourier modes and filters out the higher-frequency modes, this is
 234 then followed by an inverse Fourier transform \mathcal{F}^{-1} . Part 5 stands for a convolutional
 235 layer that is used to capture the higher-frequency modes that are missed by Part 2-
 236 4. Both AFNO outputs \boldsymbol{v}_L and reconstructed U are functions, and their derivatives
 237 $D\boldsymbol{v}_L$ and DU can be computed at any query points x and z . The solutions are
 238 constrained by the Navier-Stokes equations that govern GW dynamics. For more
 239 information, refer to the corresponding texts.

240 The CAMNet architecture is shown in **Figure 1**. The CAMNet is currently trained on
 241 horizontal wind U , so CAMNet’s input is $U(x, z, t)$. The CAMNet architecture
 242 consists of 7 parts. Below, we present a detailed computational implementation of
 243 each part. Similar to the iterative update strategy of each Fourier layer in FNO (Li
 244 et al., 2020), the improved iterative update strategy of each Fourier layer in CAMNet
 245 can be expressed as follows:

$$246 \quad \boldsymbol{v}_1 = \overset{\text{Part 1}}{\tilde{\mathcal{P}}} U(x, z, t) \quad (6)$$

$$247 \quad \boldsymbol{v}_{l+1} = \sigma \left(\overset{\text{part 5}}{\tilde{\mathcal{C}}} \boldsymbol{v}_l + \overset{\text{part 2}}{\mathcal{F}^{-1}} \overset{\text{part 3}}{\tilde{\mathcal{R}}} \overset{\text{part 4}}{\mathcal{F}} \boldsymbol{v}_l \right), \text{ for } l = 1, 2, \dots, L \quad (7)$$

$$248 \quad U(x, z, t + 1) = \overset{\text{part 6}}{\tilde{\mathcal{Q}}} \boldsymbol{v}_L \quad (8)$$

249 where \mathcal{P} and \mathcal{Q} are encoder and decoder that are realized by two neural networks that
 250 projects $U(x, z, t)$ to hidden representation \boldsymbol{v}_l and projects the representation back to
 251 the solution $U(x, z, t + 1)$. σ is a nonlinear activation function. The additional term \mathcal{C}
 252 is a convolutional layer that acts on \boldsymbol{v}_l . \boldsymbol{v}_l denotes the output of the l -th Fourier layer
 253 of AFNO. \mathcal{F} denotes the Fourier transform that acts on \boldsymbol{v}_l , \mathcal{R} is a linear transform
 254 layer that acts on $\mathcal{F}(\boldsymbol{v}_l)$ to handle its low-frequency modes. \mathcal{F}^{-1} is an inverse Fourier
 255 transform that acts on $\mathcal{R}(\mathcal{F}(\boldsymbol{v}_l))$. The details of each part are provided below:

256 **Part 1: Encoder \mathcal{P}** The encoder is implemented using the token embedding layer in
 257 the Vision Transformer architecture proposed by Dosovitskiy et al. (2020). This layer
 258 applies a linear projection to each patch to obtain a fixed-sized vector, which is then
 259 concatenated with positional embeddings representing the spatial location of the
 260 patch. The resulting sequence of vectors serves as the input to the sequent neural
 261 network layers.

262 **Part 2-4: Fourier Transform \mathcal{F} , Linear Transform \mathcal{R} , and Inverse Fourier Transform**
 263 **\mathcal{F}^{-1}**

264 As stated in Li et al., 2020, since the inputs and outputs of partial differential
 265 equations (PDEs, such as Navier-Stokes equations) are continuous functions, it is
 266 more efficient to represent them in Fourier space and perform global convolution.
 267 This is due to the quasi-linear computational complexity and global properties of
 268 Fourier transform, making it a more efficient approach. The convolution in the spatial
 269 domain is equivalent to the pointwise multiplication in the Fourier domain. To
 270 capture global features in input data, a Fourier transform is first applied to the inputs,
 271 followed by a Linear Transform \mathcal{R} that acts on the lower-frequency Fourier modes by
 272 assigning weights to them. These weights will be updated during the training. Finally,
 273 an inverse Fourier transform is performed to obtain the output.

274 **Part 5: High-frequency Information Compensation \mathcal{C}** In each Fourier layer, we utilize
 275 a convolutional layer to extract high-frequency information because it can amplify

276 high-frequency components. Therefore, we train a convolutional layer \mathcal{C} on the
 277 outputs of Part 1 to extract their high-frequency information. As a complement to
 278 Parts 2-4, the convolutional layer enables the forward prediction of high-frequency
 279 information.

280 **Part 6: Decoder \mathcal{Q}** Given two future states independently predicted by Parts 2-4 and
 281 Part 5, we combine them and train a non-linear decoder using a multi-Layer
 282 perceptron layer with a tanh activation function to transform the AFNO outputs back
 283 into U .

284 **Part 7: Physics Informed Loss \mathcal{L}_p**

285 As CAMNet is currently trained on horizontal wind U , we utilize only the momentum
 286 flux equation (1) for the physics-informed part, assuming a constant density ρ over
 287 time. Thus, equation (1) can be simplified to

$$288 \quad \frac{\partial(\rho U)}{\partial x} + \frac{\partial(\rho W)}{\partial z} = 0 \quad (9)$$

289 In equation (9), the vertical wind W is obtained from CGCAM simulations at each
 290 prediction time step, and ρ is set to its initial values. The only physics-informed
 291 variable to be calculated in CAMNet is the derivative of U with respect to x . To enable
 292 multi-variable predictions in the future, the physics-informed part will need to
 293 involve the complete Navier-Stokes equations.

294 The loss function of CAMNet for optimizing equations (6), (7), and (8) is defined as

$$295 \quad \mathcal{L} = \alpha \mathcal{L}_{data} + \beta \mathcal{L}_p = \alpha |U - \hat{U}| + \beta \left| \frac{\partial(\rho U)}{\partial x} + \frac{\partial(\rho W)}{\partial z} \right| \quad (10)$$

296 where α and β control the weights of data-driven and physics-informed part in loss
 297 functions, respectively.

298 **Table 1:** Key model parameters used in CAMNet model and training.

Hyperparameter	Value
Batch Size	32
Learning rate	5×10^{-4}
Learning rate schedule	Cosine
Patch size (Case 1/2)	$10 \times 10/8 \times 8$
Number of AFNO layers	8
Heads number	8
Heads depth	6
AFNO embedding dimension	768
Activation function	GeLU
Dropout	0

299 CAMNet is highly optimized so that it can be trained efficiently on massively parallel
300 GPU resources. The initial application of CAMNet has shown that CAMNet can
301 achieve order-of-magnitude speedup over numerical model CGCAM. We performed
302 two cases to evaluate the performance of CAMNet (see “Result” section), and the
303 relevant model parameters are presented in **Table 1**.

304 3. Results

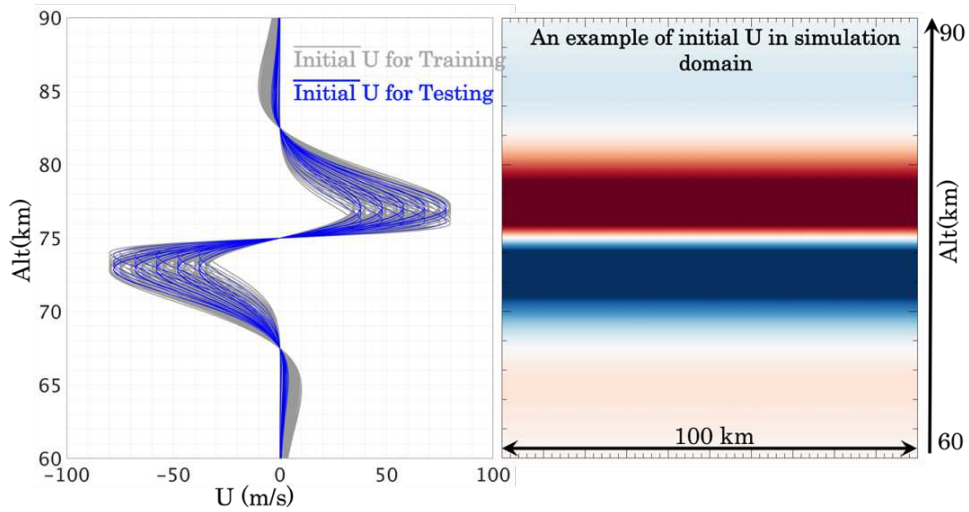
305 This section describes our initial efforts using CAMNet to accelerate simulations of
306 GWs arising from two very different sources constrained to a two-dimensional (2-D)
307 domain. Case 1 describes the generation of initial GWs by large-scale, shear-induced
308 Kelvin-Helmholtz Instability (KHI) and the successive generation of SGWs at much
309 larger scales that readily propagate to much higher altitudes. Case 2 describes
310 mountain waves arising from flow over idealized terrain, their attainment of large
311 amplitudes, breaking, and generation of SGWs that likewise attain very high
312 altitudes. CAMNet wind fields and spectra are compared with high-resolution 2-D
313 CGCAM simulations in both cases.

314 3.1 Case 1: Gravity Waves emitted from Kelvin–Helmholtz Instability

315 We explore here the potential of CAMNet for modeling GWs emitted from Kelvin-
 316 Helmholtz Instability (KHI) described by Dong et al.(2023). We use CGCAM to
 317 generate training and testing data for CAMNet. The initial background winds (see
 318 Figure 2) are specified as

$$319 \quad U(z) = U_0 \cos \left[\frac{\pi(z - z_0)}{15 \text{ km}} \right] \tanh \left(\frac{z - z_0}{h} \right) \quad (11)$$

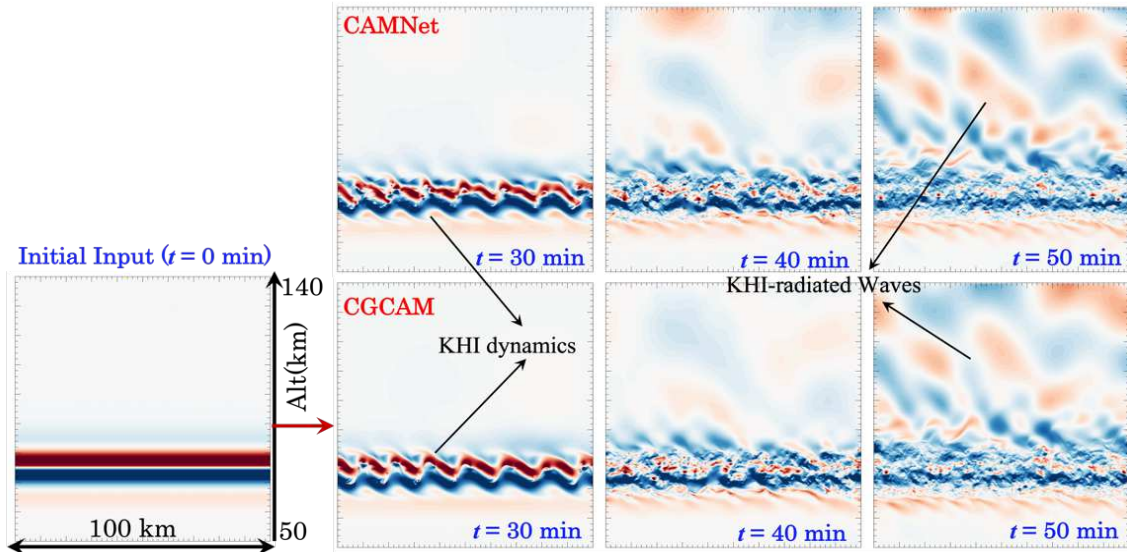
320 CGCAM simulations are performed for a computational domain having dimensions
 321 $180 \text{ km} \times 180 \text{ km}$ (x, z) with resolutions of 50 m at the shear center, with exponential
 322 mesh stretching approaching the upper and lower boundaries to reduce
 323 computational demands. Periodic boundary conditions are used at the lateral
 324 boundaries. An isothermal no-stress wall condition is used at the lower boundary,
 325 and a characteristic radiation condition is used at the upper boundary. The vertical
 326 boundary conditions are supplemented with sponge layers having 20-km depths to
 327 further ensure no reflected GWs. After excluding irrelevant data in the sponge layers,
 328 the variable U are stored on a grid of dimension of 2000×1000 .



329

330 **Figure 2:** Initial conditions for generating training and testing data (left) and an
 331 example of initial U in the simulation domain (right) for the KHI cases.

332 Given the initial conditions, CAMNet is required to simulate the future states of
 333 variable U at $t \in \{1,2,3, \dots, 50\}$ min for a suite of initial conditions. A total of 200 cases
 334 are generated by varying U_0 and h in Equation 11, and the corresponding outputs of
 335 CGCAM serves as the true reference solutions for each case. The CGCAM simulations
 336 for each case were run for 50 minutes at an interval of 1 minute. The 200 CGCAM
 337 cases were then split into a training set of 180 cases and a testing set of 20 cases. The
 338 training set is used to train CAMNet, and the testing set is used to evaluate the
 339 model's performance. Keeping the testing set separate from the training set is crucial
 340 to obtain an unbiased estimate of the model performance. All samples have a grid of
 341 2000×1000 . The CAMNet training is implemented in a multi-GPU environment
 342 with 4 V100 GPUs. Convergence is observed after approximately 250 epochs during
 343 the training process.



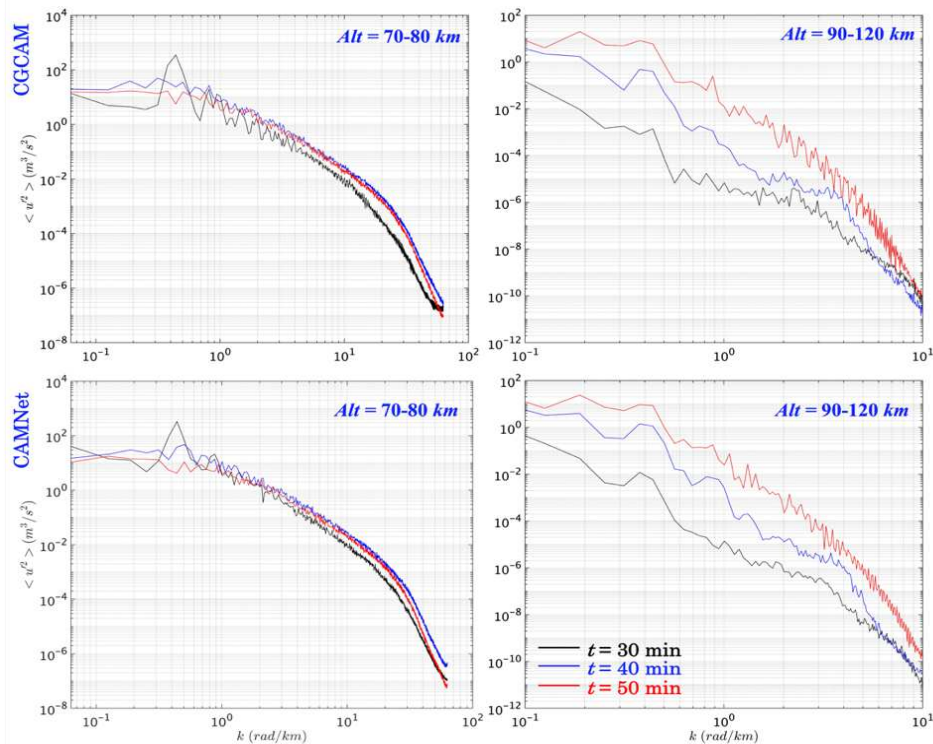
344

345

Figure 3. KHI and GWs predicted by CGCAM and CAMNet.

346 Figure 3 displays the variable U employed for both the CAMNet and CGCAM
 347 simulations during the model test, and we observe a high level of consistency between
 348 two models. The ML model CAMNet can capture small- and large-scale structures
 349 qualitatively, with clear evidence of KHI and KHI-radiated GWs seen in both
 350 CGCAM and CAMNet. The initial strong shear produces deep and broad KH billows

351 that break down after ~ 30 min, leading to the emergence of small-amplitude GWs
 352 above the KHI altitude. By 50 min, increasing GWs and apparent SGWs are seen to
 353 propagate with group velocities along the iso-phase lines that extend away from the
 354 KHI shear layer, causing GWs to propagate to higher altitudes and achieve larger
 355 amplitudes. Both models suggest that GWs are continuously emitted from the KHI
 356 and turbulence dynamics, with well-defined spatial structures having orientations
 357 and spatial scales that agree closely between the CGCAM and CAMNet fields.



358

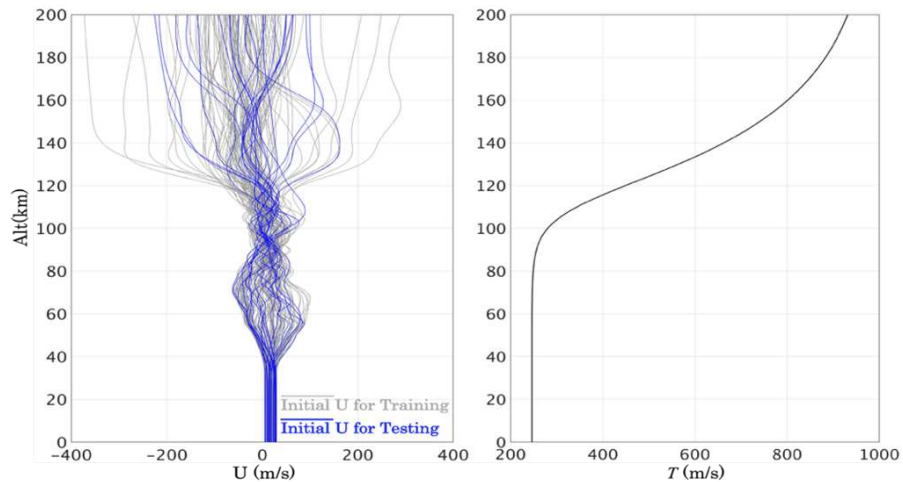
359 **Figure 4.** Case 1 spectral characteristics predicted by CGCAM (top) and CAMNet
 360 (bottom) at 70-80 km (left) and 90-120 km (right) altitude ranges at 30, 40, and 50
 361 min model simulation times (see time labels at lower right).

362 The spectral properties of GWs can reveal important details about their sources, such
 363 as the altitude and vertical extent of the source region, and the dominant
 364 wavelengths and frequencies of the GWs generated. This information can be used to
 365 improve parameterizations of GW sources in global climate models, which are critical

366 for predicting the response of the atmosphere to changing climate conditions. The
 367 spectral structures of horizontal wind disturbance u' are calculated and found to
 368 exhibit high consistency between CGCAM and CAMNet. The spectra of u'^2 computed
 369 at the KHI region and higher altitudes at $t = 30, 40,$ and 50 min are displayed in
 370 Figure 4. The spectra reflect the characteristics of KHI dynamics. In the KHI region,
 371 the onset of strong KHI and 2-D turbulence yields spectral slopes approaching $-5/3$
 372 corresponding to wavenumbers of $\sim 1-6$ rad/km. Spectral amplitudes fall sharply
 373 beyond wavenumbers of ~ 20 rad/km and exhibit steeper slopes approaching -7 in the
 374 viscous range. At higher altitudes, the $-5/3$ slope corresponds to wavenumbers of $\sim 0.8-$
 375 3 rad/km and -7 slopes correspond to wavenumbers of $3-10$ rad/km. These spectra
 376 suggest that small-scale structures discussed above are well resolved at these times.

377 These results suggest that CAMNet has promising potential as an alternative to
 378 CGCAM for simulating KHI and KHI-radiated GWs. Notably, the time cost of a single
 379 KHI case simulation using CAMNet was approximately 0.8 seconds on a single A100
 380 (80GB). This represents a significant acceleration (by > 2000) compared with the 30
 381 minutes needed by CGCAM when using 36 CPU cores.

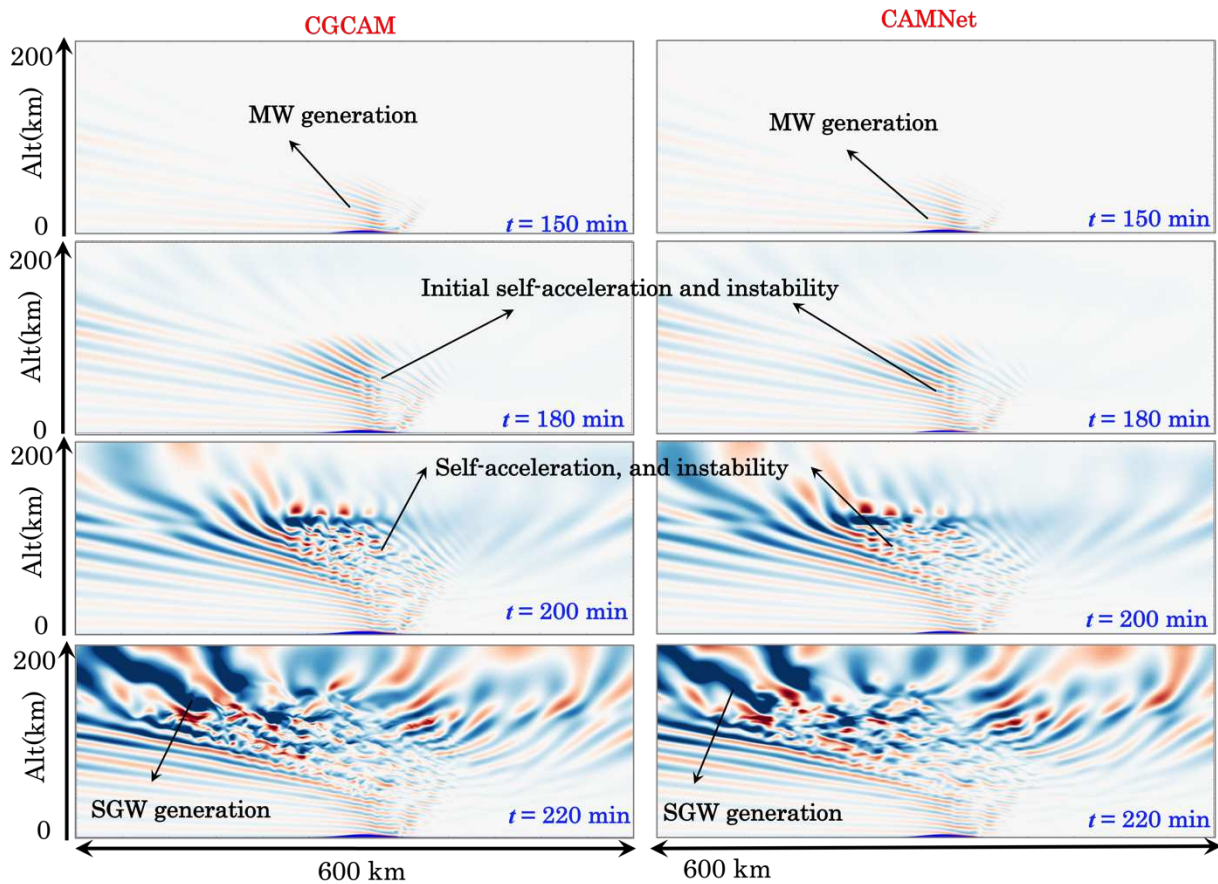
382 3.2 Case 2: Mountain Waves Generation, Propagation, and Breaking



383

384 **Figure 5.** Initial mean winds U (left) and temperature T (right) as the inputs of
 385 CGCAM for generating training and testing data for MW cases.

386 Our intent in Case 2 was to explore CAMNet capabilities for modeling MW generation,
 387 propagation, breaking, and their radiation of SGWs. MW breaking is one of the
 388 strongest sources of SGWs (Lund et al., 2020). As in Case 1, CAMNet training
 389 employs CGCAM simulation data. The CGCAM simulations cover a computational
 390 domain extending 700×220 km (x and z) at a resolution of 1 and 0.5 km in x and z .
 391 As in Case 1, the lateral boundary is periodic. At the lower boundary, a Gaussian
 392 terrain of peak height 4 km and half-width of 30 km is used, and a characteristic
 393 radiation condition is used at the upper boundary. Sponge layers of 20 km and 50 km
 394 are added to the vertical and lateral boundaries, respectively, to ensure absorption of
 395 outgoing GWs. The variable U are stored on a grid having dimensions of 600×400
 396 after irrelevant values in the sponge layers are excluded.



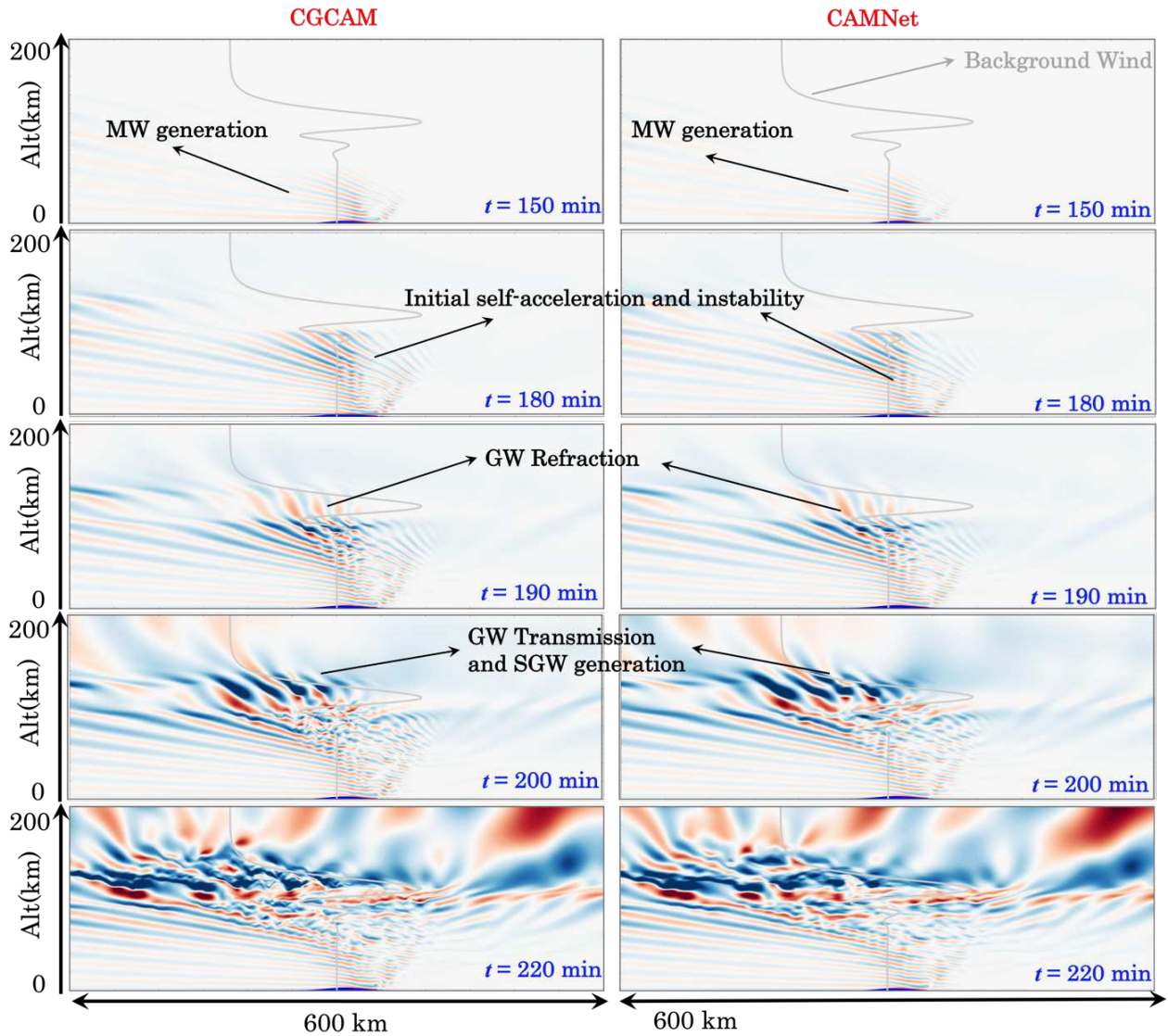
397

398 **Figure 6.** MW evolutions predicted by CGCAM (left) and CAMNet (right). The
 399 initial condition is a horizontal wind of 25 m/s at the surface flowing over a
 400 Gaussian Mountain with a height of 4 km and a half width of 30 km (see Figure 4).

401 Given the samples of initial conditions, CAMNet is trained to reproduce the future
 402 states of variable U at $t \in \{140, 145, 150, \dots, 220\}$ min. Note that we start from $t = 140$
 403 mins to avoid CAMNet being trained with non-physical data produced by CGCAM at
 404 early simulation times. A total of 200 cases are generated by varying the initial wind
 405 field, with a random wind field randomly extracted from the HWM14 at $30^\circ S, 70^\circ W$
 406 (Andes Lidar Observatory) at 00:00 on 200 days among 365 days. The initial
 407 temperature field is simplified as used in Dong et al. (2020). Winds at lower altitudes
 408 from HWM14 are consistently lower than actual observations, and thus a correction
 409 is needed to enable simulation of MW generation. To account for this discrepancy, we
 410 randomly assign wind values ranging from 0-30 m/s at these lower altitudes to
 411 facilitate the occurrence of MWs. The initial fields were assumed to be uniform over
 412 the domain. The initial wind and temperature fields are shown in Figure 5. The
 413 corresponding output of CGCAM serves as the true reference solution for each case.
 414 CGCAM simulations for each case were run for 220 minutes at an interval of 5
 415 minutes. The CGCAM simulations were then split into a training set of 180 cases and
 416 a testing set of 20 cases. All samples have a grid of 600×400 . The training of CAMNet
 417 is implemented in a multi-GPU environment with 4 V100. Convergence is observed
 418 after approximately 320 epochs during training process.

419 The CAMNet model achieves excellent skill in modeling MW generation, propagation,
 420 breaking, and SGW generation. As an illustrative example, we choose a case with a
 421 uniform initial background wind of 25 m/s. We begin with an overview of the major
 422 features of the MW and SGW evolution from $t = 150$ -220 min. Figure 6 shows U at
 423 $t = 150, 180, 200, 220$ min generated by CAMNet and the corresponding CGCAM
 424 results at these times. Considering CAMNet results first, the earliest responses at
 425 $t = 150$ min reveal MW generation at lower altitudes and their extension into the
 426 MLT. At $t = 180$ min, initial SA dynamics and instabilities are seen at lower altitudes.

427 At $t = 200$ min, there is evidence for strong SGW excitation in the MW breaking
 428 regions. The MW field and its associated instabilities and SGWs continue to intensify
 429 to $t = 200$ min. The CAMNet results approximate the CGCAM ground truth
 430 remarkably well over 220 min. Additionally, high consistency is found between
 431 CGCAM and CAMNet in the u' spectra, which are shown in the first and second rows
 432 of Figure 8, respectively.

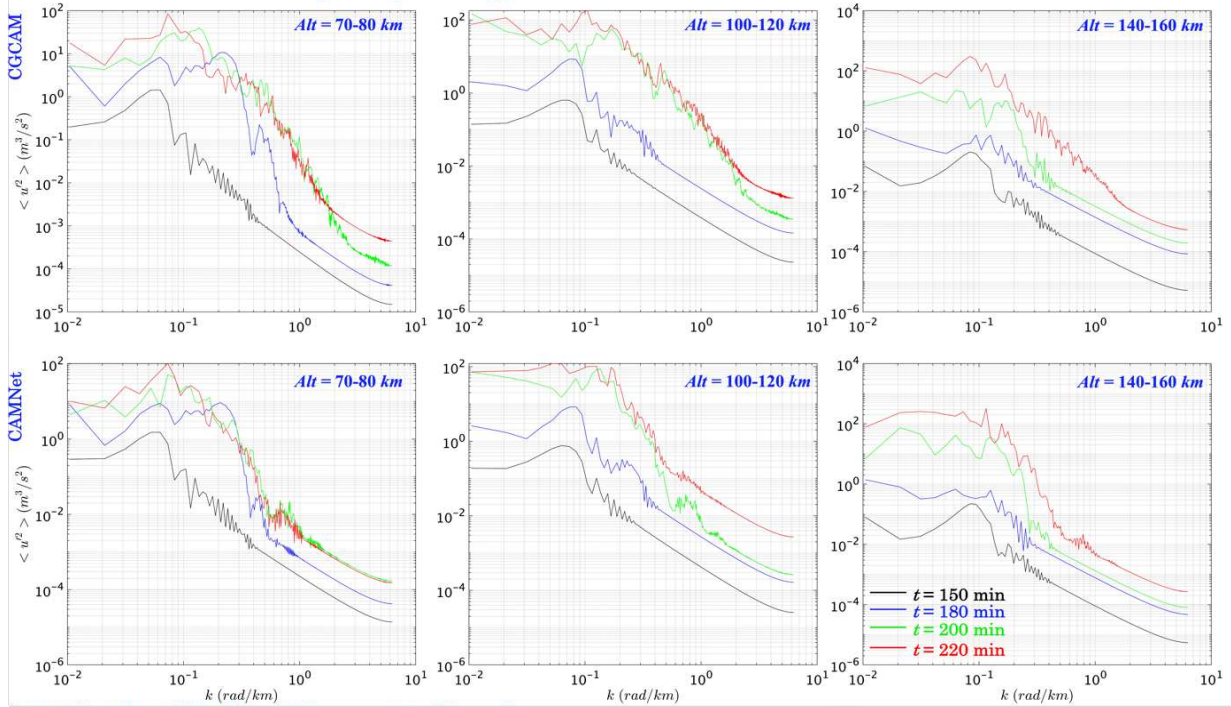


433

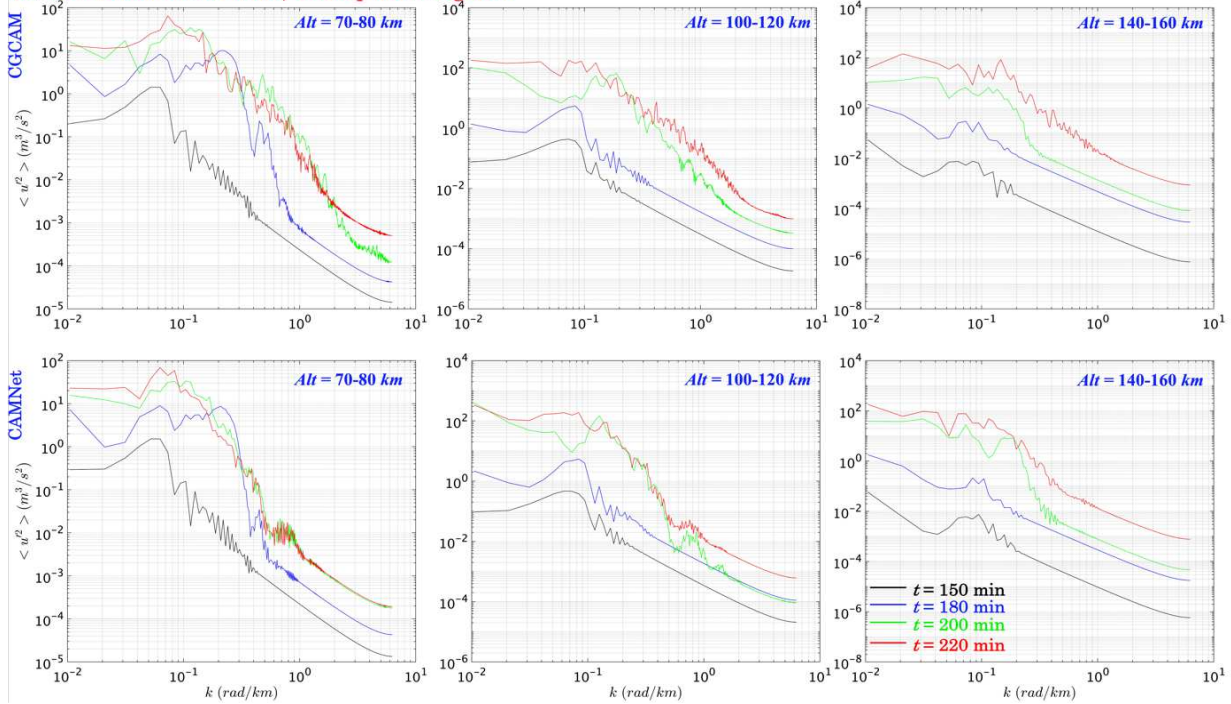
434

Figure 7. Same as Figure 6, but for a tidal wind background.

Initial Condition: Constant Wind, Correspond to Figure 5



Initial Condition: Tidal Wind, Correspond to Figure 6



435

436 **Figure 8.** Case 2 spectral characteristics predicted by CGCAM (first and third rows)
 437 and CAMNet (second and fourth rows) at 70-80 km (left), 100-120 km (middle) and
 438 90-120 km (right) at 150,180,200, and 220 min (see time labels at lower right).

439 As a further test of the generalization ability of the trained CAMNet model, we chose
440 an initial wind field that includes a tidal wind field (represented by grey lines in
441 Figure 7) that was not part of the 200 cases used for training and testing. CAMNet
442 results were compared to those of CGCAM, and also showed high consistency in
443 capturing the MW dynamics (see Figures 7 and 8), including the following:

- 444 1. Major MW responses, including MW generation, propagation, breaking, local
445 instabilities and dissipation, and SGW generation;
- 446 2. strong MW breaking, instabilities, and 2-D turbulence dynamics;
- 447 3. SGWs modulated by tidal winds and having large scales and large influences
448 extending into the thermosphere. Responses include refraction by tidal winds,
449 and reduced SGW responses at higher altitudes relative to the case of no tidal
450 wind (as shown in Figures 5 and 6); and
- 451 4. CAMNet exhibits highly consistent MW spectral characteristics with CGCAM
452 (see the third and fourth rows in Figure 7).

453 This case demonstrates that CAMNet exhibits good generalization abilities.

454 CAMNet demonstrates promising potential as a competitive alternative to CGCAM
455 for simulating MW generation, propagation, and breaking. A single MW case
456 simulation using CAMNet takes approximately 0.5 seconds on a single A100 (80GB),
457 in this case ~4000 times faster than the corresponding CGCAM simulation using 36
458 CPU cores, which took around 40 minutes.

459 4. Discussion

460 GWs play significant roles in the transport of energy and pseudo-momentum through
461 Earth's atmosphere. However, current numerical weather prediction and climate
462 models lack the resolution needed to describe the smaller-scale GWs and SGWs
463 accounting for the large majority of GW energy and pseudo-momentum fluxes to
464 higher altitudes. Thus, parameterizations are employed to represent unresolved GW
465 influences in most global atmospheric models. These typically have three primary

466 components: (1) specification of GWs at source levels, (2) GWs propagation with
467 altitude, and (3) GW dissipation and further parameterized body forcing and mixing.

468 The primary function of GW parameterizations as currently applied in global models
469 is to compute the wave-driven force on the mean flow, where the mean flow refers to
470 the grid-box mean, and the waves are intended to represent sub-grid unresolved GW
471 anomalies. These parameterizations treat GWs as linear, hydrostatic, vertically
472 propagating waves in a steady ambient environment with Boussinesq governing
473 equations. However, these assumptions and simplifications limit the representation
474 of many observed GW characteristics, which are physically well understood (e.g.,
475 Eyring et al., 2006, Hertzog et al., 2012; Eckermann et al., 2015; Stephan et al., 2016).
476 Some ML-based GW parameterization schemes have been proposed to enhance the
477 accuracy of GW parameterization (e.g., Chantry et al., 2021; Espinosa et al., 2022).
478 These ML-based GW parameterization schemes rely on traditional GW
479 parameterizations and hence are limited by the assumptions and simplifications
480 inherent in them. CAMNet offers several advantages over traditional and previous
481 ML-based GW parameterization schemes. The training process of CAMNet does not
482 rely on any existing GW parameterization schemes, thus it is not limited by their
483 assumptions and simplifications. CAMNet is trained with high-resolution simulation
484 data from the state-of-the-art atmospheric model CGCAM, which are accurate
485 numerical solutions of the Navier-Stokes equations. Well-trained CAMNet is capable
486 of resolving multi-scale and highly nonlinear GW dynamics, such as instability and
487 turbulence dynamics, self-acceleration, GW breaking, and SGW generations, at much
488 faster speed. To the best of our knowledge, CAMNet is the first ML-based approach
489 that can directly simulate highly nonlinear GW dynamics.

490 To further enhance the application of CAMNet's exceptional GW simulation
491 capabilities, CAMNet will be optimized and extended 1) from single-variable
492 simulation to multi-variable simulation to better capture the complex nonlinear
493 interactions among various GW variables; and 2) from 2D to 3D to more accurately

494 model the horizontal evolution characteristics of GWs, which are only fully displayed
495 in 3D cases.

496 5. Summary

497 In this paper, we developed a machine learning model solving the compressible
498 Navier-Stokes equations in our Complex Geometry Compressible Atmosphere Model
499 (CGCAM) named CAMNet. CAMNet is a hybrid machine learning model that
500 combines data-driven and physics-informed approaches. It is based on the Adaptive
501 Fourier Neural Operator (AFNO) proposed by Guibas et al., 2021, with modifications
502 tailored to our simulations. The main improvements include: 1) the addition of
503 convolutional layer branch to compensate for high-frequency components truncated
504 by the Fourier layers, making the model more robust in resolving multi-scale
505 dynamics, and 2) the incorporation of physical information from the Navier-Stokes
506 equation in CGCAM. The CAMNet feedback neural network utilizes a loss function
507 that combines both the physical information in Navier-Stokes equations and the data
508 loss from CGCAM simulations. This approach reduces the need for extensive training
509 data and improves the model generalization ability.

510 We evaluated the performance of CAMNet with two test cases: the first one described
511 the KHI and the associated GW radiation as explored in Dong et al., 2023, and the
512 second one addressed the generation, propagation, and breaking of MWs, which are
513 one of the most important GW components. Both cases involve small-scale instability
514 and turbulence dynamics, as well as larger-scale GWs. Our main findings from the
515 simulations are as follows:

- 516 1) CAMNet shows excellent skill on simulating the formation and intensification
517 of KHI, and KHI-radiated GWs. Those results qualitatively match the ground
518 truth remarkably well over a period of 50 mins.
- 519 2) CAMNet has excellent skill on the simulations of MW generation, propagation,
520 and breaking. CAMNet captures the major MW responses, including MW

521 generation, propagation into MLT, strong MW breaking via vortex ring
522 formation, strong instability and turbulence dynamics, intense SGW
523 generation, and strong modulation by tidal wind.

524 3) CAMNet can be trained on high-resolution CGCAM simulations under various
525 conditions and has the potential to significantly accelerate GW simulations
526 while maintaining high accuracy.

527 Accounting for the GW forcing in global atmospheric models is challenging due to
528 their limited model resolutions. A well-trained CAMNet can produce the simulations
529 orders of magnitude faster than CGCAM without any noticeable accuracy loss. This
530 has two important implications. Firstly, high-resolution GW simulations can be
531 generated within seconds, thus enabling estimation of well-calibrated and
532 constrained uncertainties regarding unresolved GW scales with higher confidence
533 compared to current global models that have severely simplified GW
534 parameterization schemes due to computational cost. Secondly, CAMNet is suitable
535 for rapidly testing hypotheses about mechanisms of GW forcing and their
536 predictability. Moreover, there is potential to develop a software library of well-
537 trained CAMNet models to be applied to a broad range of conditions. The well-trained
538 models have the potential to become a viable alternative to current GW
539 parameterizations in global models.

540 **Data availability**

541 Outputs from model simulations used in this study are archived on the ERAU High
542 Performance Computing System Vega and can be made available upon publication.

543 **Code availability**

544 The CGCAM code utilized in this study is proprietary and can only be accessed
545 through collaboration. The original AFNO code is available at
546 <https://github.com/lonestar686/AdaptiveFourierNeuralOperator>.

547 **Acknowledgement**

548 Research described here was supported by the Air Force Office of Scientific Research
 549 (AFOSR) grant FA9550-18-1-0009 and NSF grants AGS-1759471, AGS-2032678,
 550 AGS-2131350, and AGS-2128443. We also acknowledge Embry-Riddle Aeronautical
 551 University and the National Center for Atmospheric Research for access to
 552 supercomputer platforms that allowed the CGCAM and CAMNet simulations
 553 reported here. The National Center for Atmospheric Research is a major facility
 554 sponsored by the National Science Foundation under Cooperative Agreement No.
 555 1852977.

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